



Array Simulation and Beamforming for the Expanded GMRT

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Giant Metrewave Radio Telescope (GMRT)

- ❑ GMRT is one of the most sensitive telescopes for studying astrophysical phenomena at low radio frequencies (50 to 1450 MHz). GMRT is a national project of the Govt. of India
- ❑ Located 80 km north of Pune, 160 km east of Mumbai
- ❑ Array telescope consisting of 30 antennas of 45 m diameter – processing through a sensitive radio receiver and real-time digital signal processing backend



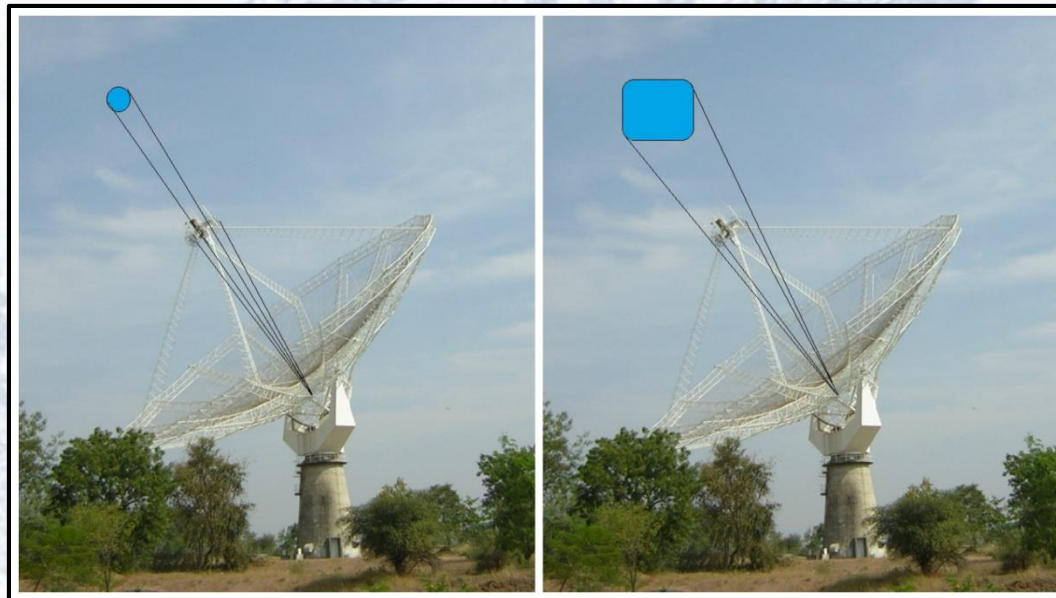
Panoramic View of the GMRT Array

The Expanded GMRT (eGMRT)

- 30 new antennas at baselines less than 5 km. : need correlator and beamformer for 30 antennas

problem statement

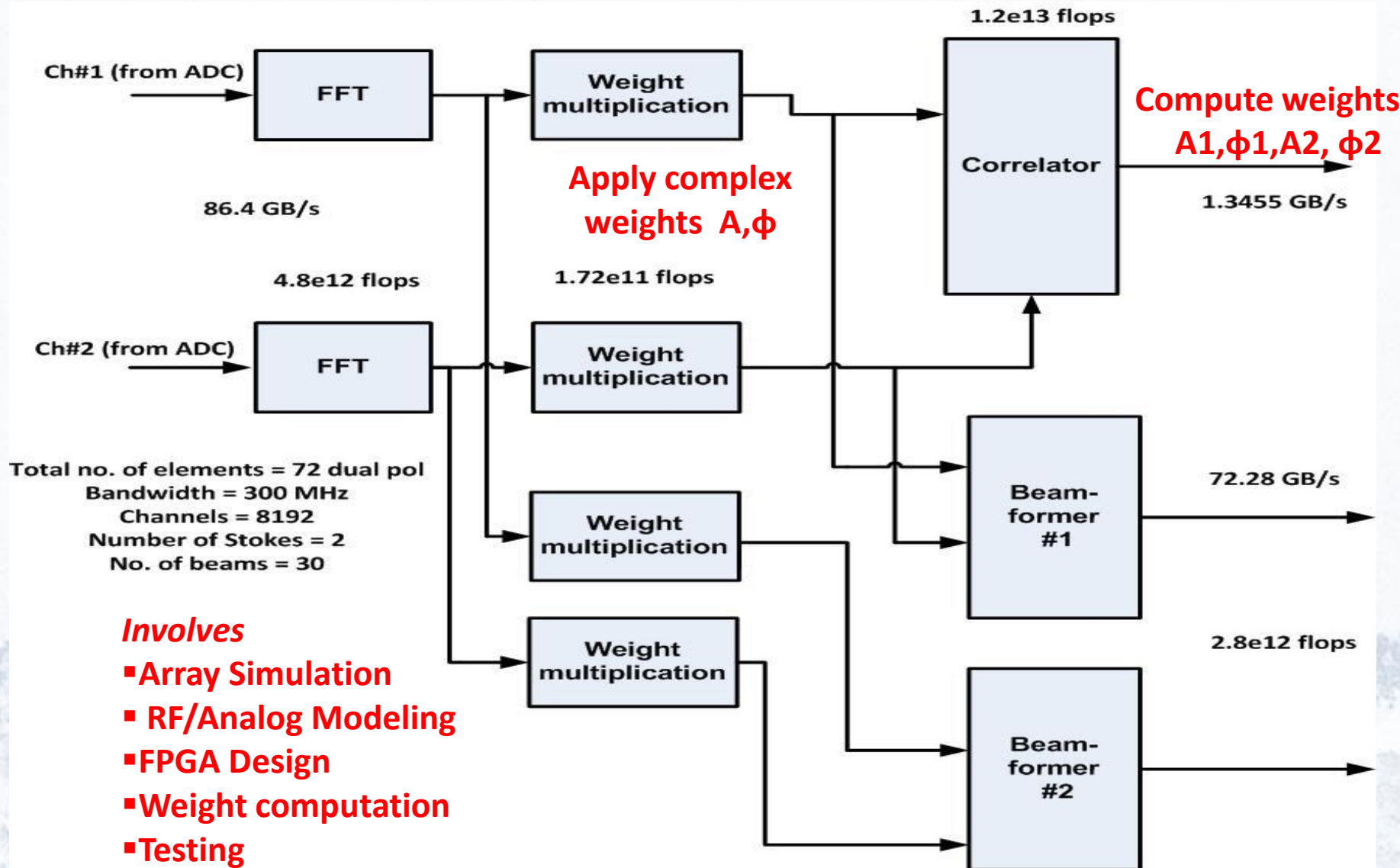
- Focal Plane Array (FPA) feeds with 30 beams on the sky : **system-level simulation, design and test multi-beam beamformer, calibration**
- 550-850 MHz RF, 300 MHz bandwidth, 16384 spectral channels, 30 beams – analog receiver and multi-beam digital beamformer



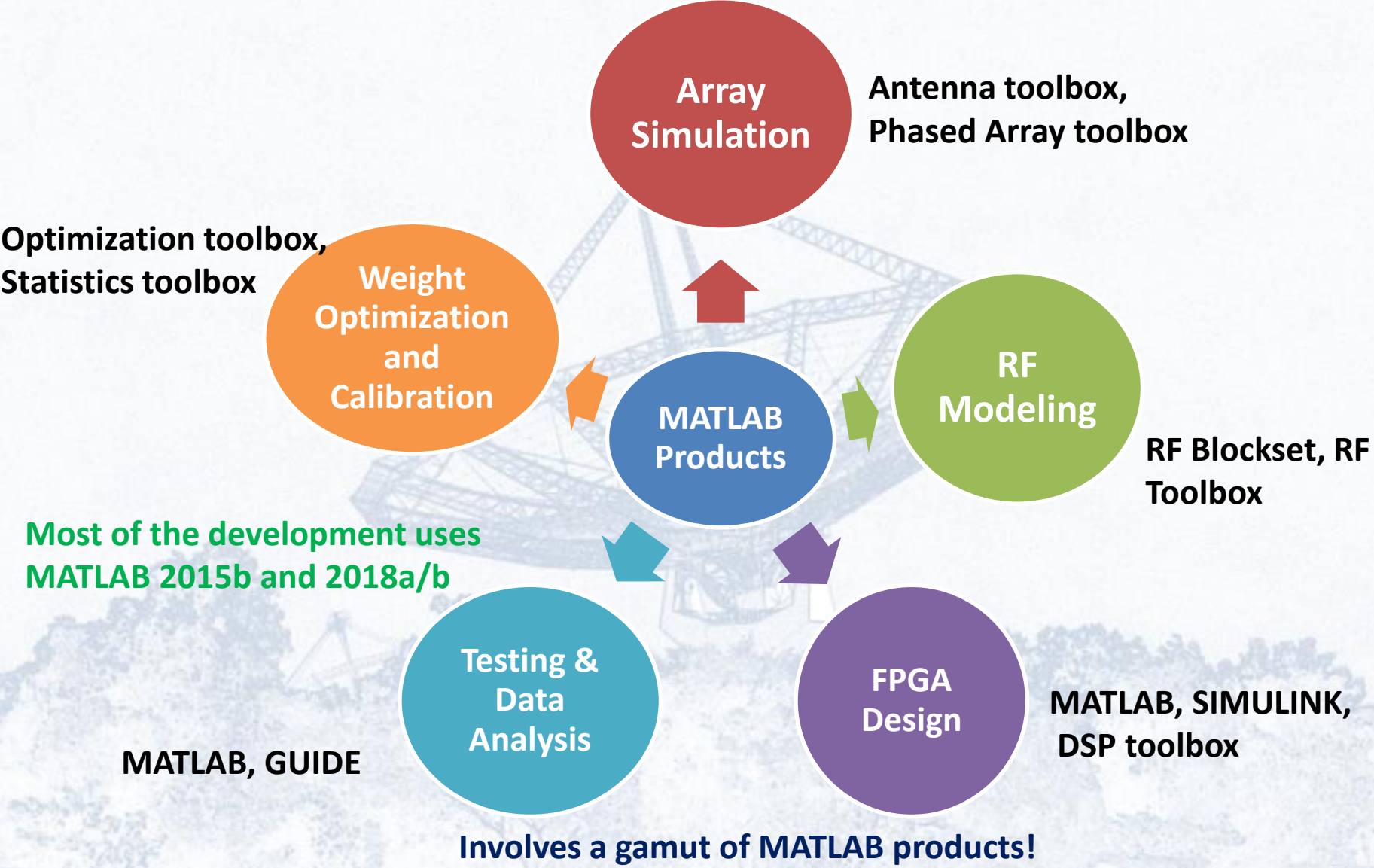
**Artist's Impression:
Increased Field-of-
View with FPA at
the focus
(not to scale)**

*Refer: The
Expanded GMRT
Patra et al.,
MNRAS, 483,
2019*

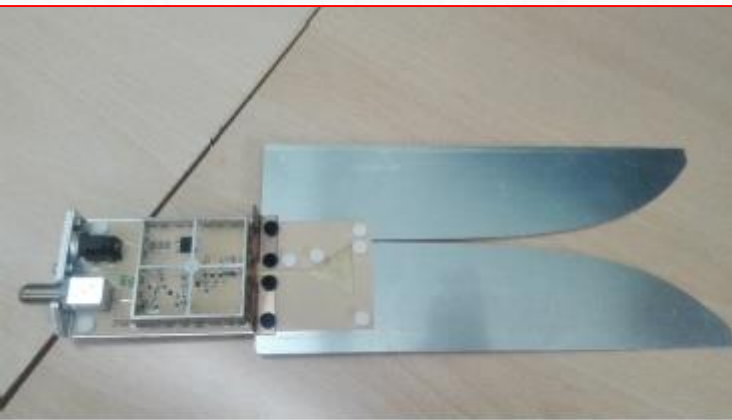
Block diagram: Basic FPA beamformer



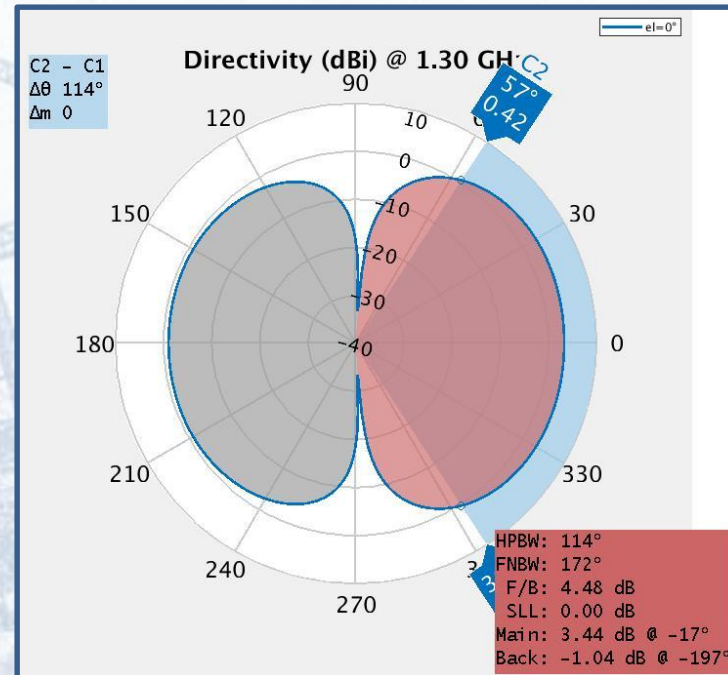
Beamformer Development using MATLAB



Simulating Vivaldi Element



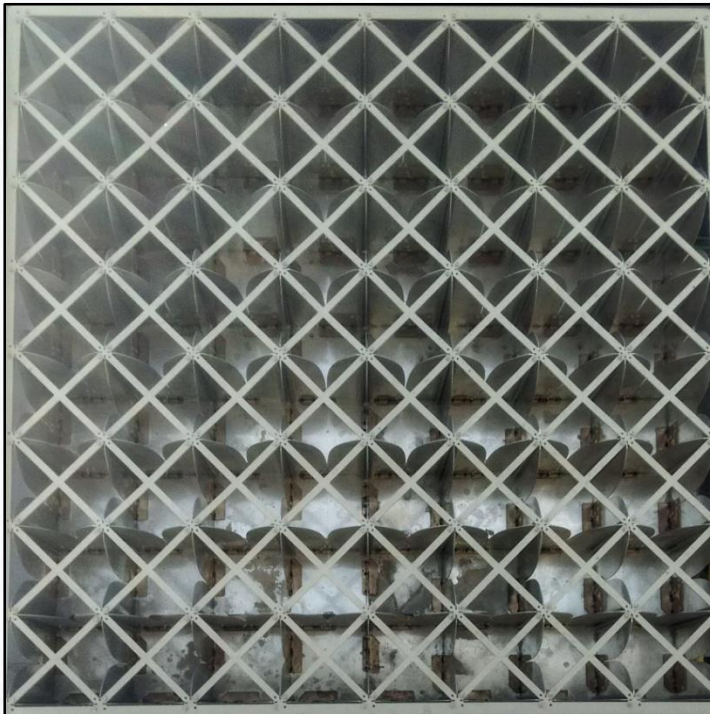
Vivaldi Antenna Element along with the Front-end Electronics



Element radiation pattern at 1.3 GHz simulated using Antenna Toolbox

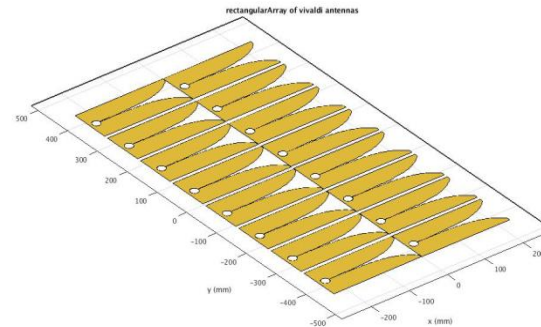
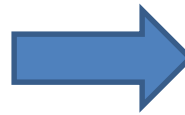
'TaperLength',179.92e-3 , 'ApertureWidth',83.94e-3
, 'SlotLineWidth',0.5e-3 , 'CavityDiameter',20e-3
, 'CavityToTaperSpacing',23.61e-3
, 'GroundPlaneLength',230.50e-3
, 'GroundPlaneWidth',100e-3, 'FeedOffset',-65.0e-3
(all dimensions in m)

Simulating Vivaldi Array

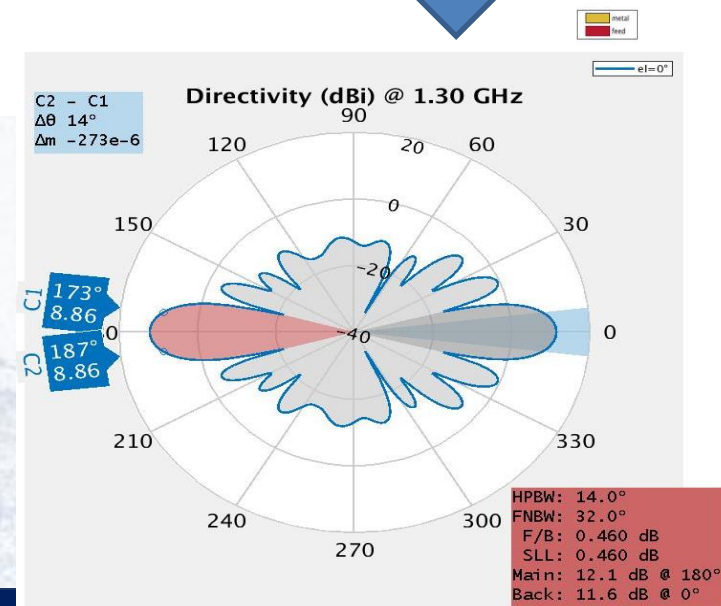


8x9 Vivaldi Array (from ASTRON) at GMRT

Two rows/columns (rectangular array) 8x2



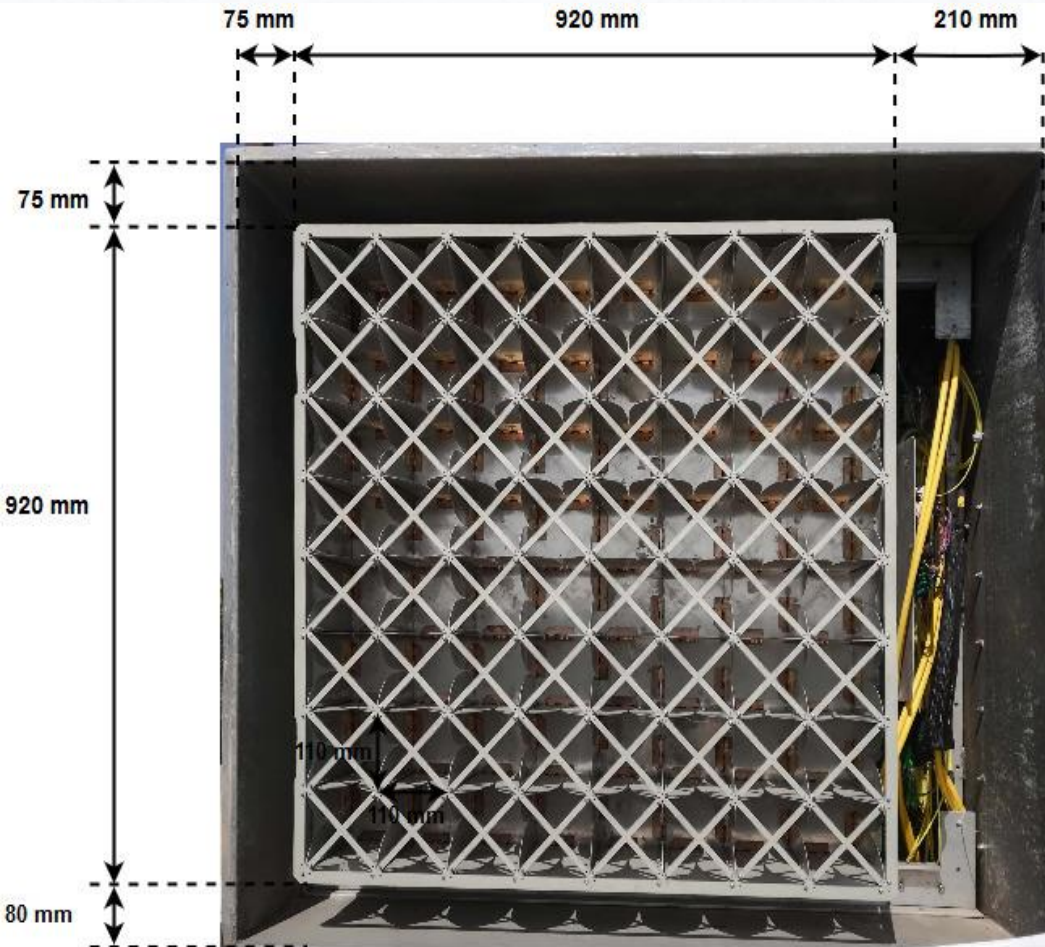
Radiation pattern



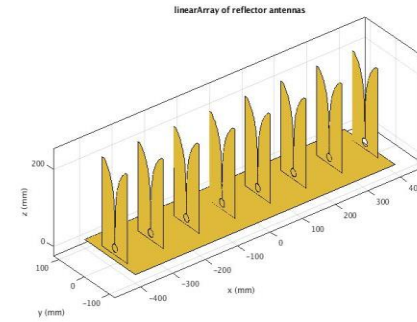
The spacing between the elements is $\sim\lambda/2$ at 1.4 GHz – elements are mutually coupled

Current aim is to understand the antenna simulation at the system-level

Simulating Vivaldi Array



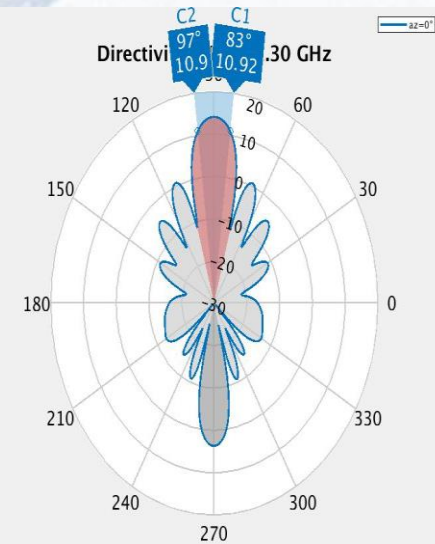
Actual Vivaldi array installation in an enclosure



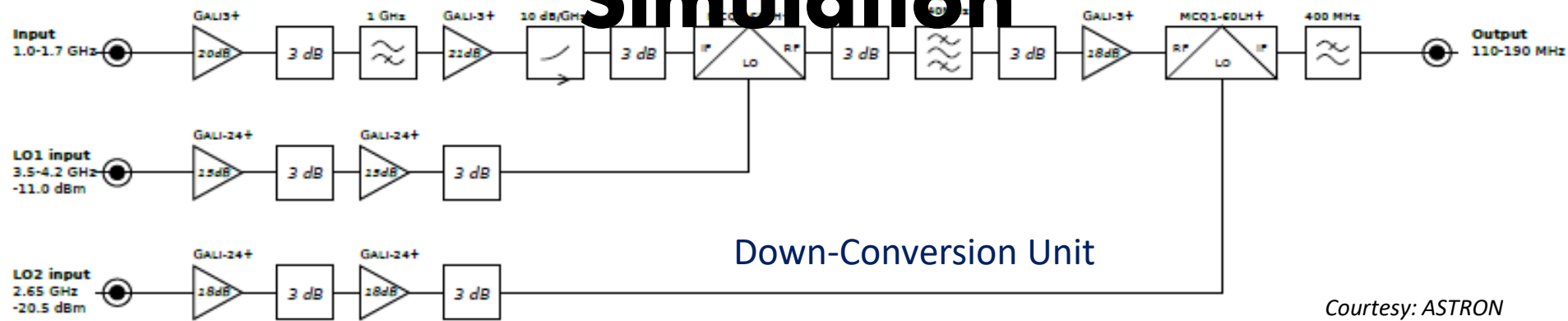
metal
feed

Typical Simulation (8x1 with ground plane)

C2 - C1
 $\Delta\theta$ 14°
 Δm -24.9e-3

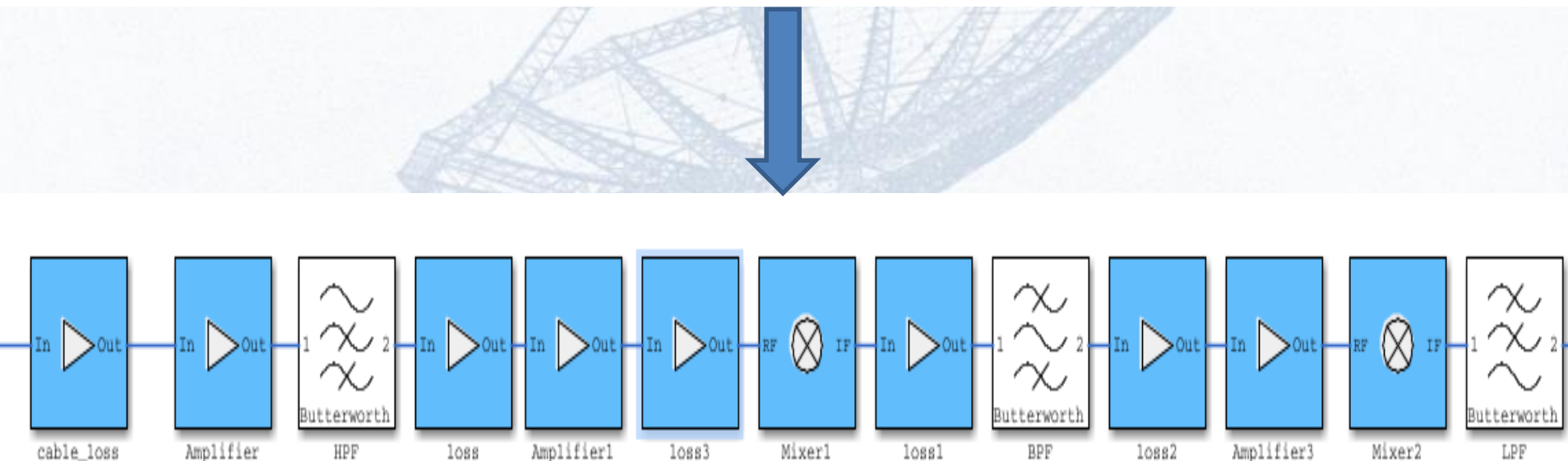


Down-Conversion Unit: Simulation



Down-Conversion Unit

Courtesy: ASTRON

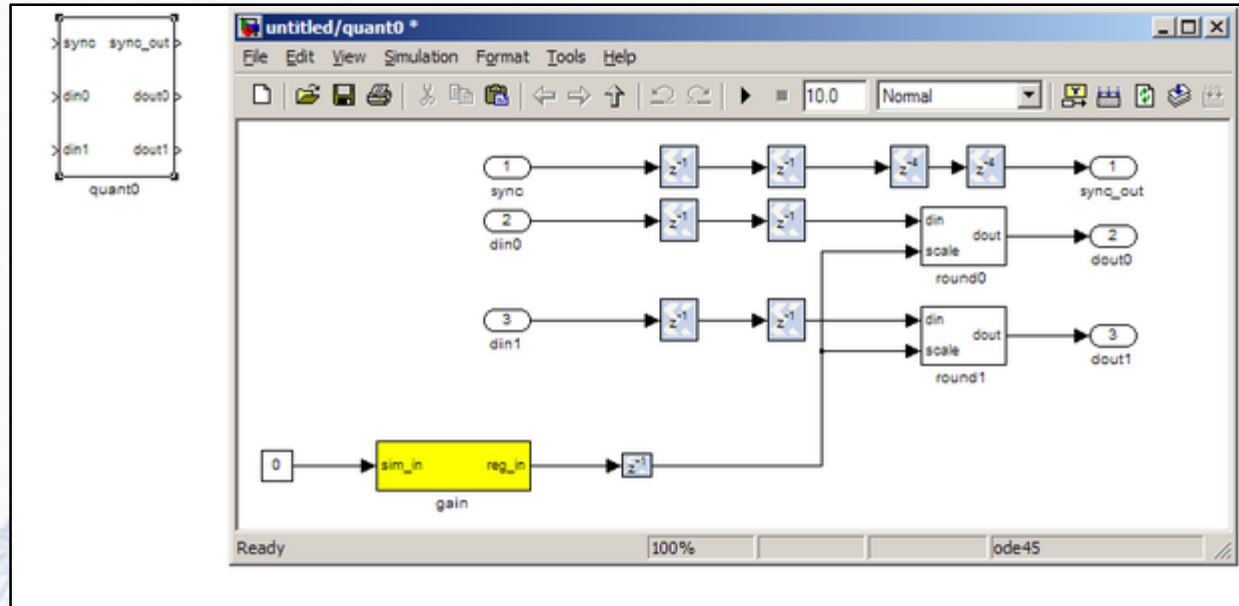


Simulation Model of DCU Block

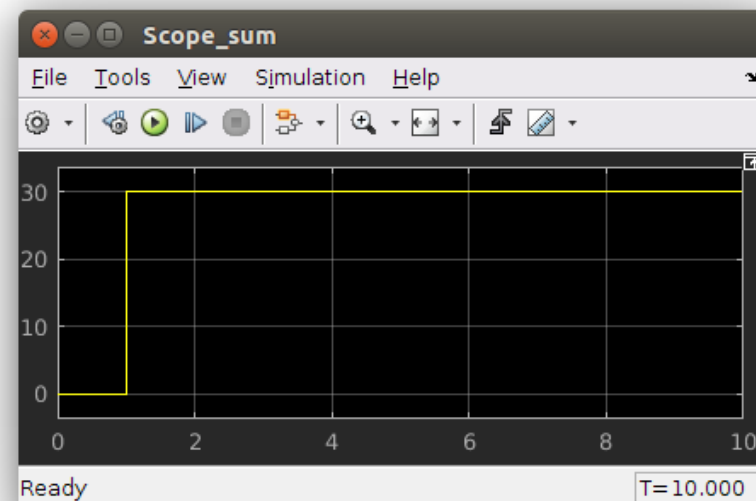
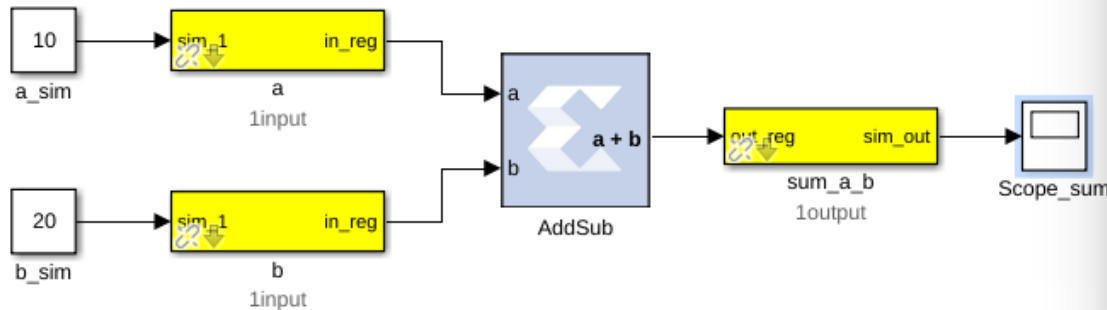
FPA Beamformer: FPGA Design

Model-based design approach

Helped in behavioral simulation ahead of implementation – leads to faster prototype development

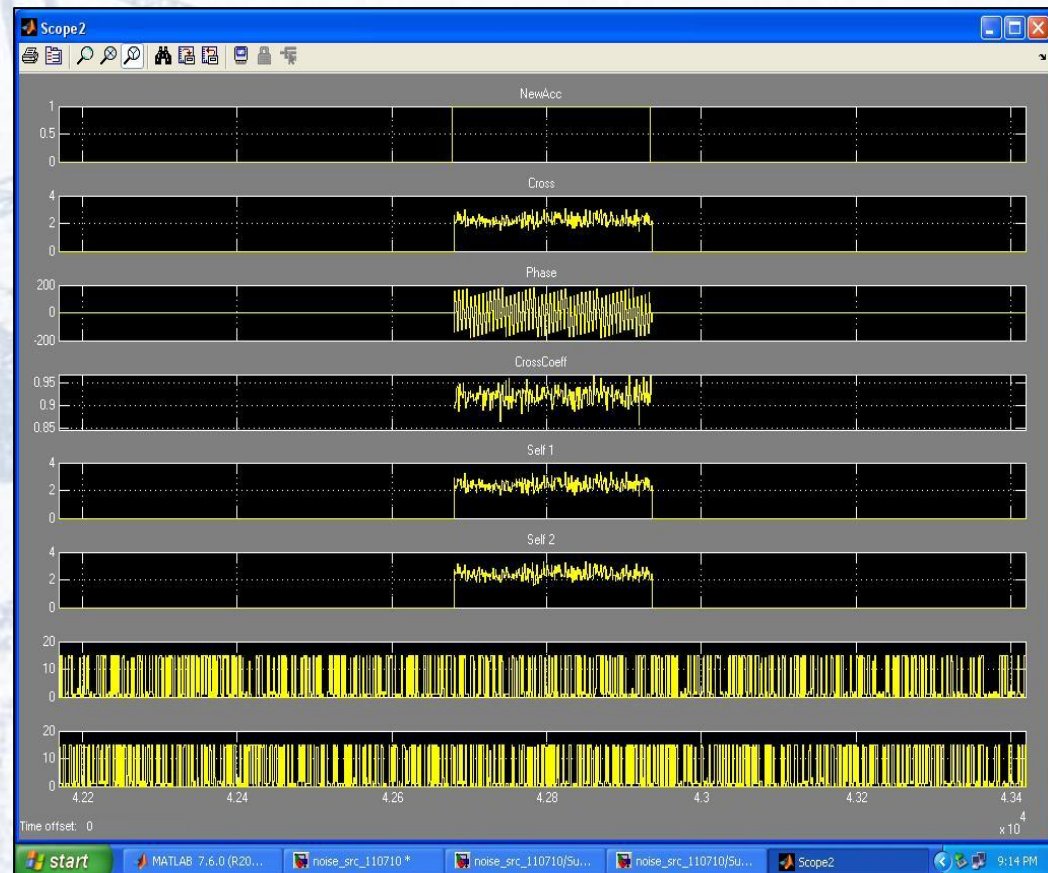
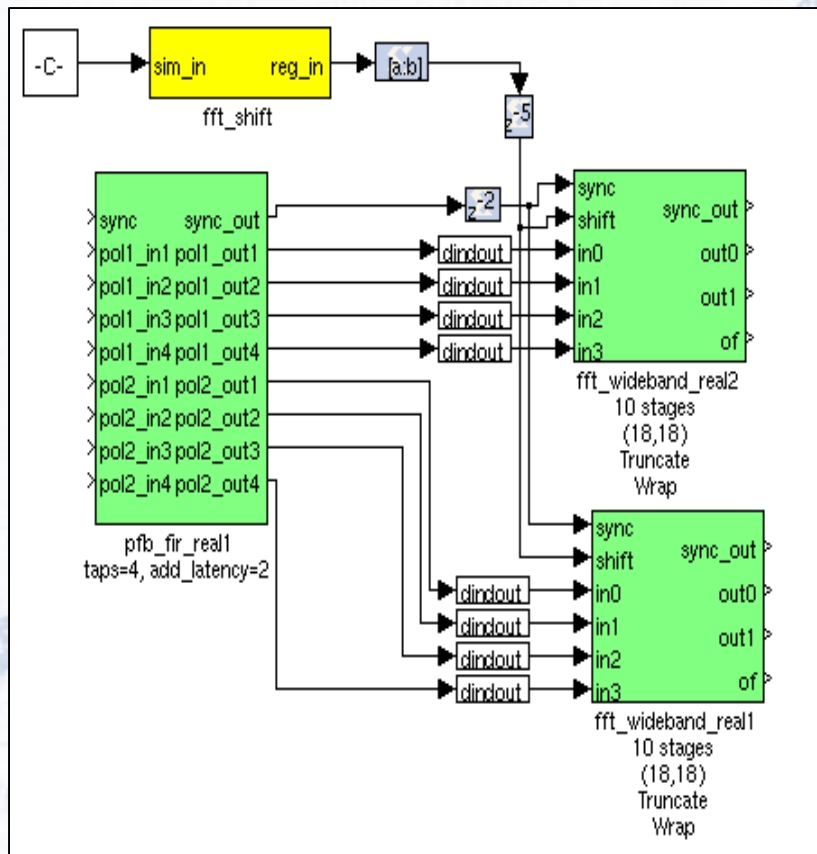


Snapshot of design



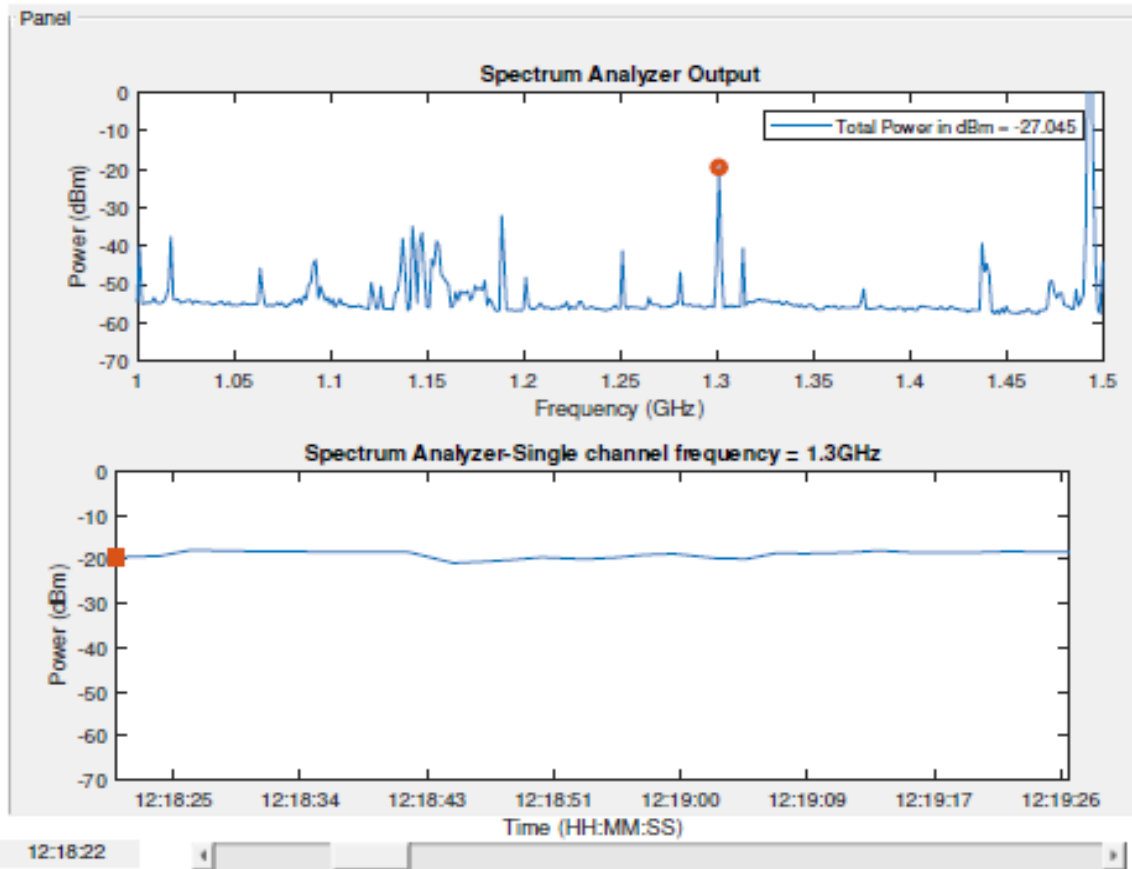
Behavioural Simulation

- ❑ Uses CASPER tool-flow (<https://casper.berkeley.edu/>)
- ❑ Ease of generating complex test scenario and test vectors



Monitoring and Diagnostic Tools

- ❑ GUIDE (GUI Development Environment) used for monitoring and recording (.avi file) the spectrum to determine the gain stability of the system
- ❑ Helps in visualizing time-varying external radio interference and its effects on the other signal quality



Import file and plot data

Start Time 2018 3 21 12 16 27

End Time 2018 3 21 12 30 47

Enter channel number

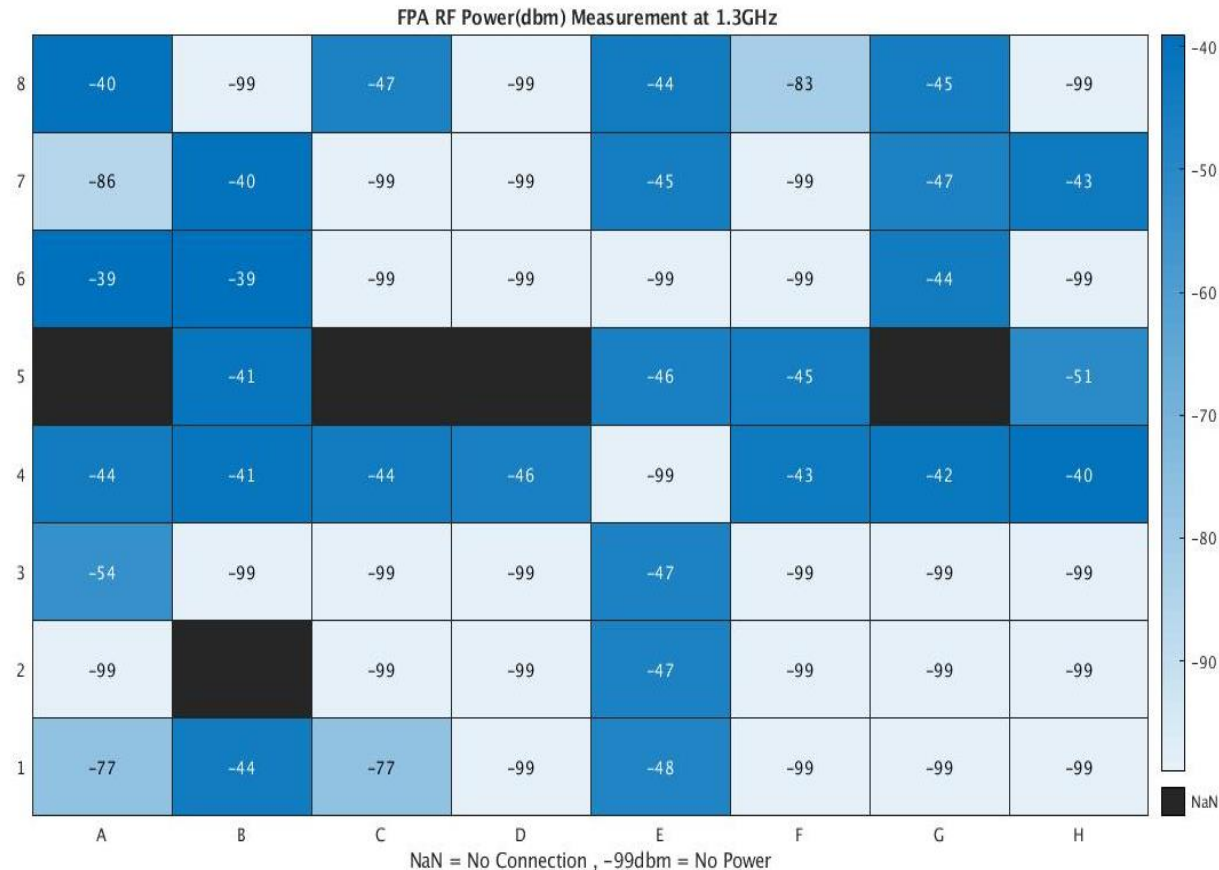
Movie_START TIME

Movie_END TIME

Snapshot of spectrum monitoring and recording GUI

Monitoring and Diagnostic Tools

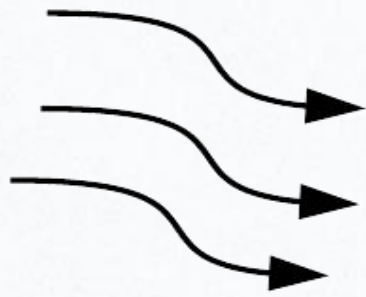
RF power measurement for choosing appropriate elements for the beamforming process



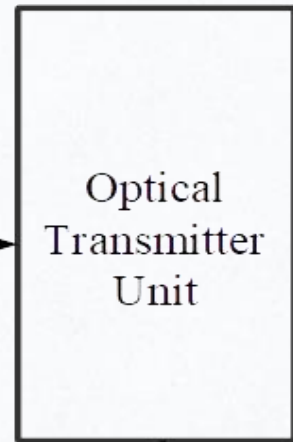
RF power of the array elements (free-space testing) –*heatmap* function



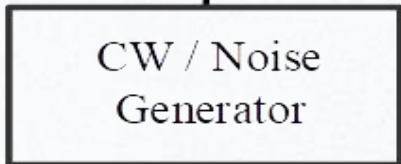
3m dish



FPA



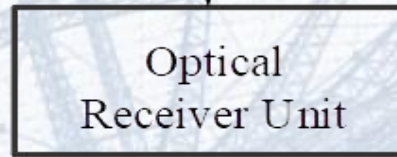
Optical Transmitter Unit



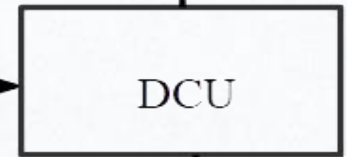
CW / Noise Generator

~67m

Free-Space Test Range



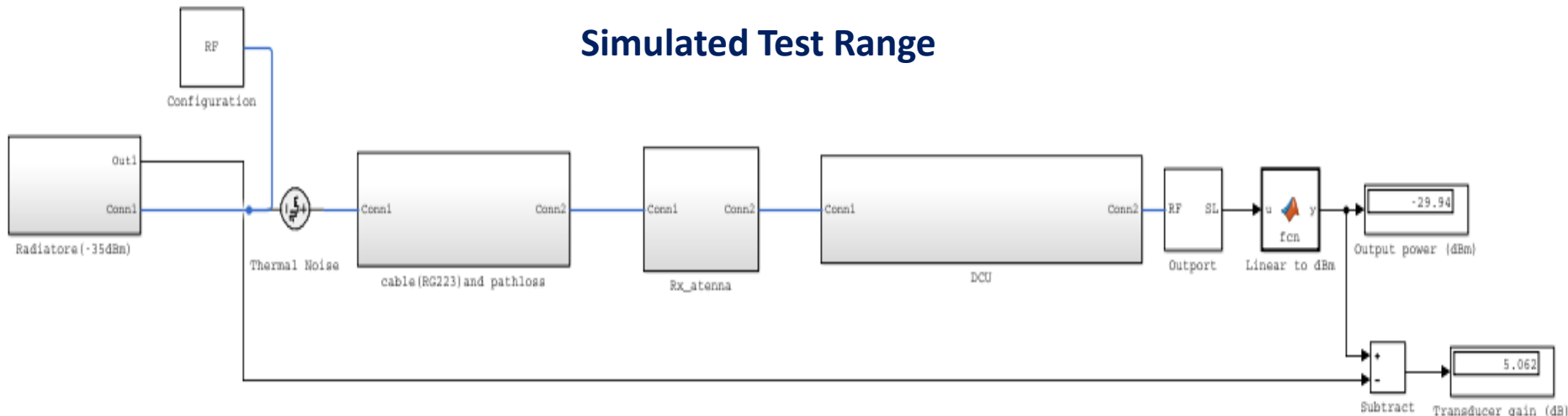
Optical Receiver Unit



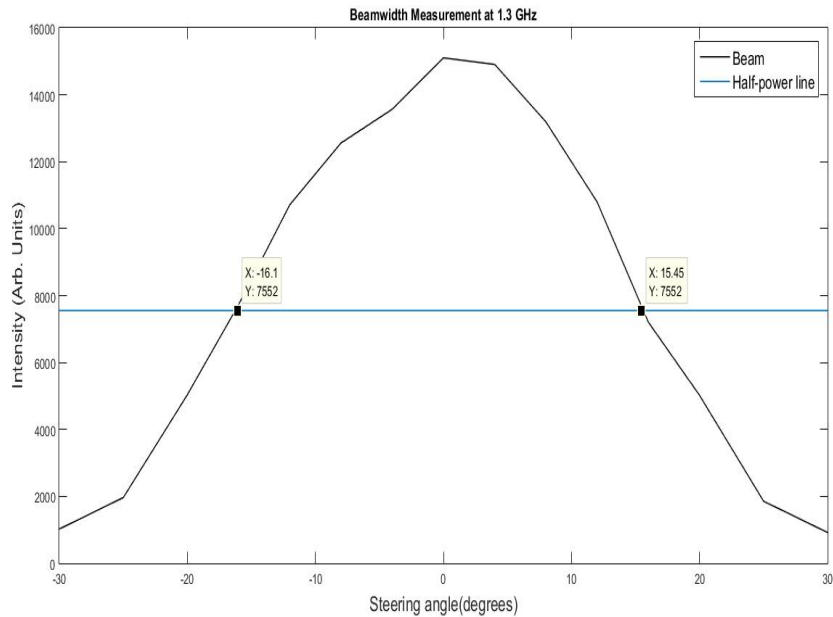
DCU

ADC

Simulated Test Range



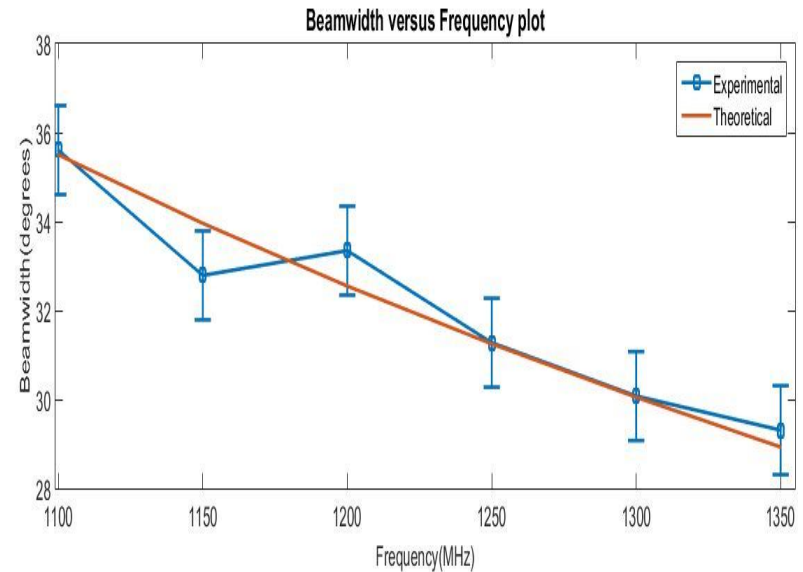
Test Results



❑ Beamsteering across radiating antenna; good match between theoretical and measured beamwidth

❑ Test carried out for a linear-array configuration (4-element array with 11 cm spacing)

❑ Comparison between theoretical beamwidth and measured beamwidth as a function of frequency



Beam optimization and Calibration

Beamformer weights optimization
(Max-SNR method):

```
[V,D] = eig(Ron-Roff, Roff);  
[tmp,idx] = max(abs(diag(D)));  
w_maxsnr = V(:,idx);
```

R_{on} and R_{off} are the on-source and off-source Array Covariance Matrices (ACM)

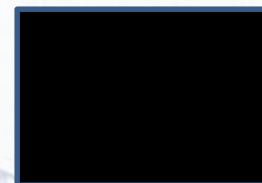
Off- source (sky background)

Compute Off-
source Array
Covariance Matrix

Celestial Radio Source



Compute On-
source Array
Covariance Matrix



~5°



ACM values
from
correlator

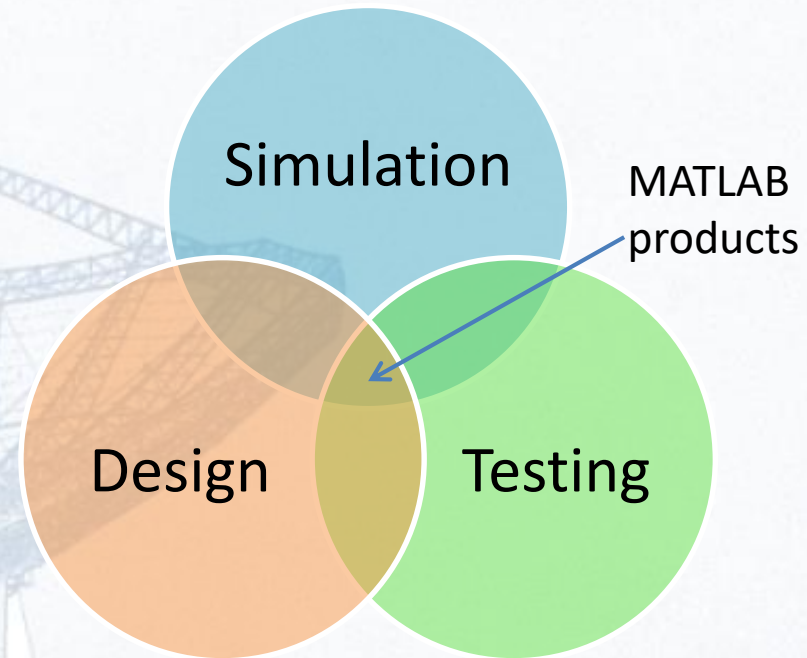
Optimum
Weight
Calculation

Applying
Weights to
the
beamformer

Summary

Beamformer Development

- ❑ Current status of array simulation and beamformer development for the Expanded GMRT was described.
- ❑ Simulation was carried out from the system-level modeling point-of-view (to understand the input to digital system)
- ❑ Several MATLAB and products were used during the beamformer development and data analysis
- ❑ Would like to learn about the upcoming features in MATLAB products



MATLAB and products form a common platform for the development

Acknowledgements

Team members (past and present) : Bela Dixit, Priya Hande, Aamer Shaikh, Rahul Argade, eGMRT beamformer development team

Short-term interns – Ritwik Sarkar, Aditi Patade

The Expanded GMRT project team

GMRT Engineering Groups

The CASPER Collaboration

ASTRON, The Netherlands

Shashank Kulkarni, Mathworks (India)

Mathworks (India)

Thank You!

